Study Guide for Unit on Space

The Universe

| Evide | nce for the above theory include: | | |
|---|--|--|--|
| o shift in light as a star moves away from a viewer | | | |
| 0 | shift in light as a st | | |
| 0 | Edwin used the red/blue shift in the light of stars to show that not | | |
| | | y from Earth, they are actually moving away from each | |
| | are la | arge clouds of gas and are where stars are born. | |
| | oc | | |
| | | form after very large stars experience a supernova. | |
| | | f a collection of stars, gas, and planets, moons, and rocks. | |
| | | is the name given to our galaxy. | |
| What | is the difference between a moon and | a planet? | |
| al Mo | tion | | |
| uo | | object's resistance to motion and depends upon the | |
| objec | t's mass and speed. | | |
| The L | aw of | states that an object in motion will stay in motion and an | |
| | | d upon by a | |
| | | s to a pulling (attraction) force between any two objects | |
| | due to the bending of space by an obje | | |
| Gravit | ty depends upon the | of an object and the | |
| | een the two objects. | | |
| As pla | nets move through space, their own | has them moving in a straight | |
| path. | However, the pull of | from the Sun pulls the planets towards the | |
| Sun. | The combination of | and | |
| | | I the Sun or each moon around a planet. | |
| rstan | ding Our Solar System | | |
| | | is the model of our solar system with Earth as the | |
| cente | r of our solar system. | | |
| | | is the model of our solar system with the Sun as the | |
| cente | r of our solar system. | | |
| | | was the first person to develop the model of our | |
| solar | system with the Sun at the center of ou | | |
| | | mathematically proved that the Sun was at the | |
| | r of solar system with the planets orbiti | _ | |
| Keple | r's 1 st Law of Planetary Motion states – | that the orbit of every planet is an | |
| with t | he Sun located at one | | |

| | is when the Earth is closer to the Sun and occurs in | |
|---|---|----|
| | January. | |
| 0 | is when the Earth is farther from the Sun and occurs in | |
| | July. | |
| Keplei | 's 2 nd Law of Planetary Motion states – Planets travel faster when they aret | .c |
| | n and when they are farther from the Sun. | |
| Keplei | 's 3 rd Law of Planetary Motion states – Planets to the Sun take less time | |
| to orb | it the Sun than planets from the Sun. | |
| | was the first person to provide visible evidence that the | |
| | located at the center of our solar system and that the planets orbit the Sun and moons orbit | |
| planet | | |
| y fron | n the Sun | |
| | is the state of most matter within our Sun and is the most | |
| comm | on state of matter in the universe. | |
| | a is not made up atoms but is instead made up of | |
| | | |
| The te | mperatures inside the Sun are so high that charged particles and small atoms collide with such | |
| | hat they fuse together to form larger, while releasing tons of | |
| | in a process called nuclear while releasing tons of | |
| Durin | nuclear reactions, large atoms split apart into smaller atoms, while | ` |
| During | ng tons of energy. Nuclear fission reactions occur in nuclear bombs and at nuclear power plant | = |
| | | |
| | ergy created in the Sun travel to earth as energy which | 1 |
| uses _ | waves to transfer energy. | 1 |
| uses _ Electro | waves to transfer energy. omagnetic waves all travel at the same through the | 1 |
| uses _ Electro vacuu | waves to transfer energy. magnetic waves all travel at the same through the m of space but vary according to wavelength and frequency. | 1 |
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| uses _ Electro vacuu anoth | waves to transfer energy. magnetic waves all travel at the same | 1 |
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| uses _ Electro vacuu anoth within | waves to transfer energy. magnetic waves all travel at the same | 1 |
| uses _ Electro vacuu anoth within | waves to transfer energy. magnetic waves all travel at the same | |
| uses _ Electro vacuu anoth within | waves to transfer energy. magnetic waves all travel at the same through the most space but vary according to wavelength and frequency. is the distance from the crest of one wave to the crest of er wave. is the time it takes for one wavelength to pass a given point one second and is measured in Hertz. waves have the longest wavelengths and the lowest frequency. waves have shorter wavelengths than radio waves but longer engths than infrared waves. waves are also known as heat waves and have shorter wavelengths | |
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| | sunlight to rearrange the atoms in carbon dioning $_{}$ sugar, $C_6H_{12}O_6$, and $C_8H_{12}O_6$ | |
|-------------------------------|--|----------------------------------|
| | 3ugur, c ₀ 111200, und Oxygen, 02, und | |
| Sunlight is captured by | pigments in g | reen parts of plants, algae. |
| phytoplankton, and cyanob | | reen parts or plants, algae, |
| | green light but absorbs and ι | uses red and blue light. |
| | | |
| | nts travel through the food chain as organism | |
| or animals that have eaten | | |
| All organisms use the proce | ss of | |
| | $_{12}O_6$, and oxygen, O_2 , into carbon dioxide, CO_2 | |
| energy. | | |
| Cellular respiration takes pl | ace in the | organelles. |
| | | |
| ns | | |
| The Earth takes | days to orbit the Sun. Every | years, we ad |
| | account for the extra 0.25 days. | |
| The Earth takes h | ours to rotate once around its axis in | |
| direction. | | |
| The counter-clockwise direc | tion is why the Sun rises in the | and sets in the |
| • | | |
| | at a 23.5 degree | angle in relation to the |
| ecliptic plane. | | |
| | parts of Earth experience different | |
| | the hemisphere is tilted aw | |
| | in the sky and the light intensity is | |
| | solstice is the first day of wi | |
| · | ere and June 21 in the southern hemisphere. | This is the |
| | day of the year. | |
| During the | | |
| sits | | · |
| The | | |
| northern hemisphere and D | ecember 21 in the southern hemisphere. Thi | s is the |
| | day of the year. | |
| During the | each hemisphere is not tilted av | way for towards the Sun, so |

the days and nights are equal in length. These occur on March 21 and September 21.

Tides and Eclipses

| Lunar is the name of our N | light. It does not produce | |
|--|--|--|
| any light of its own. | | |
| The Moon takes | days to orbit Earth. | |
| | phase of the Moon when it is be | etween the Sun and the Earth. |
| | phase of the Moon when it is or | n the other side of the Earth from the |
| Sun. | | |
| | phases of the Moon occur wher | n the Moon is perpendicular to the |
| Sun. | | |
| | eclipses occur when the moon | 's shadow falls on the Earth and only |
| occurs during a | moon. | |
| | vs a path across Earth called the path | າ of |
| and one can only see a so | lar eclipse when one is standing with | nin that path. |
| In order to see a total sola | ar eclipse, one must be standing in th | ne |
| shadow of the moon. | | |
| In order to see a partial so | olar eclipse, once must be standing in | n the |
| shadow of the moon. | | |
| | eclipses occur when Ea | rth's shadow falls on the Moon and |
| only occurs during a | moon. | |
| Tides are produced by the | e gravitational pull of the | on the Earth. |
| | art in Earth's tides, than the Sun, be | |
| | to Earth than the Sun. | |
| | tides occur during | and |
| r | moons, when the gravitational pull o | of the Sun is combined with the |
| gravitational pull from the really large. | Moon. During these tides, the diffe | erence between high and low tides is |
| | _ tides occur during quarter moons, | when the gravitational pull from the |
| Sun cancels some of the n during these times is not t | noon's gravitational pull. The differenthat different. | ence between high and low tides |